

APPENDIX A

TOTAL FLUX DENSITY DISTRIBUTION FROM A HELIOSTAT

Flux distribution on cylindrical surface by multiple aimed points. The following algorithm explains the calculation involved in MATLAB language.

```
clear
% Define window on cylindrical surface
theta = [0:360]; % define dimension of theta (0 - 360 degree)
L_theta = length(theta);
height = [800:1:1200]; % define dimension of tower height
L_height = length(height);
Flux = zeros(L_height, L_theta);
tower_H = 10; % tower height 10 m above the ground
radius = 2; % radius of cylinder
B = 0.3; A = 0.2; %

% Calculate sun position
day = 1; month = 3; LT = 10; % input day , month, time
Ls = -105; %Longitude Of Thailand
Lloc = -100.53; %Longitude of Bangkok, Thailand
La = 13.735; %Latitude of Bangkok
La = La*pi/180;
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```
i_d = day;
j_m = month;
if      j_m == 1
        dn = i_d;
elseif j_m == 2
        dn = 31+i_d;
elseif j_m == 3
        dn = 59+i_d ;
elseif j_m == 4
        dn = 90+i_d;
elseif j_m == 5
        dn = 120+i_d;
elseif j_m == 6
        dn = 151+i_d;
elseif j_m == 7
        dn = 181+i_d;
elseif j_m == 8
        dn = 212+i_d;
elseif j_m==9
        dn = 243+i_d;
elseif j_m == 10
        dn = 273+i_d;
elseif j_m == 11
        dn = 304+i_d;
else
        dn = 335+i_d;
end
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day_angle = 2*pi*(dn-1)/365;

T = day_angle;

Et=229.18*(0.000075+0.001868*cos(T)-0.032077*sin(T)-0.014615*cos(2*T)-
0.04089*sin(2*T))/60;

ST=LT+(4*(Ls-Lloc)/60)+(Et);

H=15*(12-ST)*pi/180;

Dec=23.45*sin(2*pi*(dn+284)/365)*2*pi/360;

Al=asin((cos(La1)*cos(Dec)*cos(H))+sin(La1)*sin(Dec)); %Altitude of sun
Az=acos((sin(Al)*sin(La1)-sin(Dec))/(cos(Al)*cos(La1))); % Azimuth of sun
if ST <=12
    A_z = Az;
else
    A_z =2*pi-Az ;
end

x = cos(Al)*cos(A_z);
y = cos(Al)*sin(A_z);
z = sin(Al);
s = [x,y,z];
J2 = sqrt((x)^2+(y)^2+(z)^2);
S = s/J2; %unit vector of sun position relative to tower base

% position of heliostats

load (heli)

L_h = size(heli,1);

for quandant = 1 : 4

    for p = 1:L_h    % calculate (Xo,Yo,Zo) for each heliostat

        if quandant == 1

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        Xo=heli(p,1);
        Yo=heli(p,2);
        Zo=heli(p,3);
elseif quandant == 2
        Xo=-heli(p,1);
        Yo=heli(p,2);
        Zo=heli(p,3);
elseif quandant == 3
        Xo=-heli(p,1);
        Yo=-heli(p,2);
        Zo=heli(p,3);
else
        Xo=heli(p,1);
        Yo=-heli(p,2);
        Zo=heli(p,3);
end

D = [ 0-Xo, 0-Yo, tower_H-Zo];
J1 = sqrt((0-Xo)^2+(0-Yo)^2+(tower_H-Zo)^2);
R = D/J1; % unit vector from mirror to aim point
Xd = R(1,1);
Yd = R(1,2);
Zd = R(1,3);

% find intercept point on cylindrical surface due to ray reflect from center of mirror by ray tracing technique

F= Xd^2+Yd^2;
G=2*Xo*Xd + 2*Yo*Yd;
K=Xo^2 +Yo^2-radius^2;

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t10= (-G + sqrt(G^2-4*F*K))/(2*F);
t20= (-G - sqrt(G^2-4*F*K))/(2*F);
    if t10 > t20
        t30 = t20;
        t40 = t10;
    else
        t30 = t10;
    end
    t40 = t20;
    if t30 > 0
        t0 = t30;
    elseif t40 > 0
        t0 =t40;
    else disp (' no t0 ')
    end

%% X0p, Y0p , Z0p are coordinate of intercept point on cylindrical surface
X0p = Xo + t0*R(1,1);
Y0p = Yo + t0*R(1,2);
Z0p = tower_H;
Dop = [X0p-Xo,Y0p-Yo,tower_H-Zo];
Jop = sqrt((X0p-Xo)^2+(Y0p-Yo)^2+(tower_H-Zo)^2);
Rs=Jop*tan(0.25*pi/180); % Radius of solar disk
Rop = Dop/Jop; % unit vector from mirror to target
mirror_azi = atan2 (Y0p, X0p); % position azimuth of center point of mirror
if mirror_azi >= 0
    mirror_azi = mirror_azi ;
else

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    mirror_azi = 2*pi+ mirror_azi;
end

% find normal vector of each heliostat
N=(S+Rop);
N = N/abs(N); %% unit normal vector of each heliostat
Alt = asin(N(1,3))*180/pi; %Altitude of mirror in degree
Azm = atan2(N(1,2),N(1,1)); % Azimuth of each mirror
if Azm >= 0
    Azm = Azm;
else
    Azm = 2*pi+Azm;
end

%% vector on mirror surface when mirror parallel with ground ( from
mirror center to four corner of mirror)
u_old1 = [A/2;B/2;0];
u_old2 = [-A/2;B/2;0];
u_old3 = [-A/2;-B/2;0];
u_old4 = [A/2;-B/2;0];

Euler1 = [cos(pi/2-Alt) 0 sin(pi/2-Alt);
          0 1 0;
          -sin(pi/2-Alt) 0 cos(pi/2-Alt) ];
Euler2 = [cos(Azm) -sin(Azm) 0;
          sin(Azm) cos(Azm) 0
          0 0 1];
Euler=Euler2*Euler1;

```

% vector on mirror surface when operate with Euler (from mirror center to four corner of mirror)

$u_new1 = Euler * u_old1$;

$u_new2 = Euler * u_old2$;

$u_new3 = Euler * u_old3$;

$u_new4 = Euler * u_old4$;

%%%% position of mirror after operate with Euler

$new_0 = [Xo;Yo;Zo]$;

$new_1 = u_new1 + [Xo;Yo;Zo]$;

$new_2 = u_new2 + [Xo;Yo;Zo]$;

$new_3 = u_new3 + [Xo;Yo;Zo]$;

$new_4 = u_new4 + [Xo;Yo;Zo]$;

% find intercept point on image plane by ray tracing

$d = -(Xod * X0p + Yod * Y0p + Zod * Z0p)$;

% point 0 (center point of mirror)

$t0 = -$

$(Xod * new_0(1,1) + Yod * new_0(2,1) + Zod * new_0(3,1) + d) / (Xod^2 + Yod^2 + Zod^2)$;

$X00p = new_0(1,1) + t0 * Rop(1,1)$;

$Y00p = new_0(2,1) + t0 * Rop(1,2)$;

$Z00p = new_0(3,1) + t0 * Rop(1,3)$;

% point 1

$t1 = -$

$(Xod * new_1(1,1) + Yod * new_1(2,1) + Zod * new_1(3,1) + d) / (Xod^2 + Yod^2 + Zod^2)$;

$X1p = new_1(1,1) + t1 * Rop(1,1)$;

$Y1p = new_1(2,1) + t1 * Rop(1,2)$;

$Z1p = new_1(3,1) + t1 * Rop(1,3)$;

```

% point 2

t2=-
(Xod*new_2(1,1)+Yod*new_2(2,1)+Zod*new_2(3,1)+d)/(Xod^2+Yod^2+Zod^2);

X2p= new_2(1,1) + t2*Rop(1,1);

Y2p= new_2(2,1) + t2*Rop(1,2);

Z2p= new_2(3,1) + t2*Rop(1,3);

% point 3

t3=-
(Xod*new_3(1,1)+Yod*new_3(2,1)+Zod*new_3(3,1)+d)/(Xod^2+Yod^2+Zod^2);

X3p= new_3(1,1) + t3*Rop(1,1);

Y3p= new_3(2,1) + t3*Rop(1,2);

Z3p= new_3(3,1) + t3*Rop(1,3);

% point 4

t4=-
(Xod*new_4(1,1)+Yod*new_4(2,1)+Zod*new_4(3,1)+d)/(Xod^2+Yod^2+Zod^2);

X4p= new_4(1,1) + t4*Rop(1,1);

Y4p= new_4(2,1) + t4*Rop(1,2);

Z4p= new_4(3,1) + t4*Rop(1,3);

%% find effective area of principal image

Area_i=1/2*abs(((X1p*Y2p)+(X2p*Y3p)+(X3p*Y4p)+(X4p*Y1p))-
((X2p*Y1p)+(X3p*Y2p)+(X4p*Y3p)+(X1p*Y4p)));

% calculate principal image size (LH,LT) and θ*

%Azm_ =Azimuth of position mirror measure from south to eastward direction
(depend on position)

%Azm_M = Azimuth of mirror (depend on time)

%Azm_S= Azimuth of sun

%A l_S= Altitude of sun

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%Al_M= Altitude of mirror (depend on time)

Azm_ = atan2(Yo,Xo);

if Azm_ >= 0

    Azm_ = Azm_;

else

    Azm_ = 2*pi+Azm_;

end

tower_al =acos ((tower_H-Zo)/Jop);

L_1=sqrt((sin(Azm_M-Azm_))^2*(cos(tower_al))^2+(cos(Azm_M-Azm_))^2);

b1=(sin((pi/2)-Al))^2+(sin(tower_al))^2-2*sin((pi/2)-
Al_S)*sin(tower_al)*cos(Azm_S-Azm_);

b2=2+2*cos(pi/2-Al_S)*cos(tower_al)-2*sin(pi/2-Al_S)*sin(tower_al)*cos(Azm_S-
Azm_);

b=sqrt(b1/b2);

L_2=sqrt((sin(Azm_M-Azm_))^2*(cos(Al_M))^2+(b*sin(tower_al)-cos(Azm_M-
Azm_)*cos(tower_al)*cos(Al_M))^2);

theta_11=asin((sin(Azm_M-Azm_)*cos(tower_al))/L_1);

theta_12=acos((cos(Azm_M-Azm_)/L_1));

sa = asin((b*sin(tower_al)-cos(Azm_M-Azm_)*cos(tower_al)*cos(Al))/L_2);

sb = acos((sin(Azm_M-Azm_)*cos(Al_M))/L_2);

theta_star = (sa-theta_12);

theta_star_dd = theta_star*180/pi;

theta_star_d = (round(abs(theta_star_dd)/5))*5;

if theta_star_d == 50

    fluxXX=flux50;

elseif theta_star_d == 55

    fluxXX=flux55;

elseif theta_star_d == 60

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```
    fluxXX=flux60;
elseif theta_star_d == 65
    fluxXX=flux65;
elseif theta_star_d == 70
    fluxXX=flux70;
elseif theta_star_d == 75
    fluxXX=flux75;
elseif theta_star_d == 80
    fluxXX=flux80;
elseif theta_star_d == 85
    fluxXX=flux85;
elseif theta_star_d == 90
    fluxXX=flux90;
elseif theta_star_d == 95
    fluxXX=flux95;
elseif theta_star_d == 100
    fluxXX=flux100;
elseif theta_star_d == 105
    fluxXX=flux105;
elseif theta_star_d == 110
    fluxXX=flux110;
elseif theta_star_d == 115
    fluxXX=flux115;
elseif theta_star_d == 120
    fluxXX=flux120;
elseif theta_star_d == 125
    fluxXX=flux125;
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elseif theta_star_d == 130
    fluxXX=flux130;
elseif theta_star_d == 135
    fluxXX=flux135;
else
    fluxXX = flux140;
end

%change coordinate
Al_ = acos (Rop(1,3));

Euler1=[cos(pi-Al_) 0 -sin(pi-Al_);
        0      1      0;
        sin(pi-Al_) 0 cos(pi-Al_)];

Euler2=[cos(Azm_) sin(Azm_) 0 ;
        -sin(Azm_) cos(Azm_) 0;
        0      0      1];

Euler=Euler1*Euler2;

point1=[X1p;Y1p;Z1p]-[X00p;Y00p;Z00p];
point2=[X2p;Y2p;Z2p]-[X00p;Y00p;Z00p];
point3=[X3p;Y3p;Z3p]-[X00p;Y00p;Z00p];
point4=[X4p;Y4p;Z4p]-[X00p;Y00p;Z00p];

Y1=Euler*point1;
Y2=Euler*point2;
Y3=Euler*point3;
Y4=Euler*point4;

Z =[Z1p Z2p Z3p Z4p];

Z = sort(Z);

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```

Y1=[Y1(1,1) Y2(1,1) Y3(1,1) Y4(1,1)];
Y2=[Y1(2,1) Y2(2,1) Y3(2,1) Y4(2,1)];
Y1=sort(Y1);
Y2=sort(Y2);
h_1=(Y1(1,1)/Rs);
k_1=(Y1(2,1)/Rs);
h_2=(Y2(1,1)/Rs);
k_2=(Y2(2,1)/Rs);
h_3=(Y3(1,1)/Rs);
k_3=(Y3(2,1)/Rs);
h_4=(Y4(1,1)/Rs);
k_4=(Y4(2,1)/Rs);
h1=(abs(sin(theta_star))/sin(theta_star))*(h_1*sin(sa)-k_1*cos(sb));
k1=(abs(sin(theta_star))/sin(theta_star))*(k_1*cos(theta_12)-h_1*sin(theta_11));
h2=(abs(sin(theta_star))/sin(theta_star))*(h_2*sin(sa)-k_2*cos(sb));
k2=(abs(sin(theta_star))/sin(theta_star))*(k_2*cos(theta_12)-h_2*sin(theta_11));
h3=(abs(sin(theta_star))/sin(theta_star))*(h_3*sin(sa)-k_3*cos(sb));
k3=(abs(sin(theta_star))/sin(theta_star))*(k_3*cos(theta_12)-h_3*sin(theta_11));
h4=(abs(sin(theta_star))/sin(theta_star))*(h_4*sin(sa)-k_4*cos(sb));
k4=(abs(sin(theta_star))/sin(theta_star))*(k_4*cos(theta_12)-h_4*sin(theta_11));
x=[h1 h2 h3 h4];
x=sort(x);
h_min=x(1,1);
h_max=x(1,4);
y=[k1 k2 k3 k4];
y=sort(y);
k_min=y(1,1);

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```

k_max=y(1,4);
Ax=h_min;
Ay=k_min;
Bx=h_max;
By=k_min;
Cx=h_max;
Cy=k_max;
Dx=h_min;
Dy=k_max;
% Move the origin of coordinate  $(\xi^*, \eta^*)$  to the corner A, calculate  $\Phi_A$ .
h1_s=Ax-Ax;
k1_s=Ay-Ay;
h2_s=Bx-Ax;
k2_s=By-Ay;
h3_s=Cx-Ax;
k3_s=Cy-Ay;
h4_s=Dx-Ax;
k4_s=Dy-Ay;
p01=0.05;
h=[h1_s-6*Rs:p01:h2_s+6*Rs];
k=[k1_s-6*Rs:p01:k3_s+6*Rs];
h=h';
l_h=length(h);
l_k=length(k);
inter_h = zeros(1,l_h);
inter_k = zeros(1,l_k);
for j = 1:l_k

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for i = 1:l_h

    if h(i) > 1 & k(j) > 1
        flA(i,j) = 1;
    elseif h(i) < -1 | k(j) < -1
        flA(i,j) = 0;
    elseif h(i) > 1 & k(j) < 1
        h(i)=1;
        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flA(i,j) =fluxXX( n+1,m+1);
    elseif h(i) < 1 & k(j) > 1
        k(j)=1;
        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flA(i,j) =fluxXX( n+1,m+1);

    else
        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flA(i,j) =fluxXX( n+1,m+1);

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        end
    end
end
% Move the origin of coordinate  $(\xi^*, \eta^*)$  to the corner  $B$ , calculate  $\Phi_B$ 
h1_s=Ax-Bx;
k1_s=Ay-By;
h2_s=Bx-Bx;
k2_s=By-By;
h3_s=Cx-Bx;
k3_s=Cy-By;
h4_s=Dx-Bx;
k4_s=Dy-By;
h=[h1_s-6*Rs;p01:h2_s+6*Rs];
k=[k1_s-6*Rs;p01:k3_s+6*Rs];
l_h=length(h);
l_k=length(k);
inter_h = zeros(l_h,1);
inter_k = zeros(1,l_k);
flB=zeros(l_h,l_k);
for j = 1:l_k
    for i = 1:l_h
        if h(i) > 1 & k(j) > 1
            flB(i,j) = 1;
        elseif h(i) < -1 | k(j) < -1
            flB(i,j) = 0;
        elseif h(i) > 1 & k(j) < 1

```

```

    h(i)=1;
    n=(1+h(i))/0.05;
    m=(1+k(j))/0.05;
    n=round(n);
    m=round(m);
    flB(i,j)=fluxXX( n+1,m+1);
elseif h(i) < 1 & k(j) > 1
    k(j)=1;
    n=(1+h(i))/0.05;
    m=(1+k(j))/0.05;
    n=round(n);
    m=round(m);
    flB(i,j)=fluxXX( n+1,m+1);
else
    n=(1+h(i))/0.05;
    m=(1+k(j))/0.05;
    n=round(n);
    m=round(m);
    flB(i,j)=fluxXX( n+1,m+1);
end
end
end

```

% Move the origin of coordinate (ξ^*, η^*) to the corner C , calculate Φ_c .

```
h1_s=Ax-Cx;
```

```
k1_s=Ay-Cy;
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```
h2_s=Bx-Cx;
```



```

k2_s=By-Cy;
h3_s=Cx-Cx;
k3_s=Cy-Cy;
h4_s=Dx-Cx;
k4_s=Dy-Cy;
k11=y(1,2)-Cy;
k12=y(1,3)-Cy;
h=[h1_s-6*Rs;p01:h2_s+6*Rs];
k=[k1_s-6*Rs;p01:k3_s+6*Rs];
l_h=length(h);
l_k=length(k);
inter_h = zeros(l_h,1);
inter_k = zeros(1,l_k);
flC=zeros(l_h,l_k);
for j = 1:l_k
    for i = 1:l_h

        if h(i) > 1 & k(j) > 1
            flC(i,j) = 1;
        elseif h(i) < -1 | k(j) < -1
            flC(i,j) = 0;
        elseif h(i) > 1 & k(j) < 1
            h(i)=1;
            n=(1+h(i))/0.05;
            m=(1+k(j))/0.05;
            n=round(n);
            m=round(m);
        end
    end
end

```

```

        flC(i,j) =fluxXX( n+1,m+1);
elseif h(i) < 1 & k(j) >1
        k(j)=1;
        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flC(i,j) =fluxXX( n+1,m+1);

else
        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flC(i,j) =fluxXX( n+1,m+1);
end
end
end

```

% Move the origin of coordinate (ξ^*, η^*) to the corner D , calculate Φ_D

```

h1_s=Ax-Dx;
k1_s=Ay-Dy;
h2_s=Bx-Dx;
k2_s=By-Dy;
h3_s=Cx-Dx;
k3_s=Cy-Dy;
h4_s=Dx-Dx;

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```

k4_s=Dy-Dy;
k11=y(1,2)-Dy;
k12=y(1,3)-Dy;
h=[h1_s-6*Rs:p01:h2_s+6*Rs];
k=[k1_s-6*Rs:p01:k3_s+6*Rs];
l_h=length(h);
l_k=length(k);
inter_h = zeros(l_h,1);
inter_k = zeros(1,l_k);
flD=zeros(l_h,l_k);
for j = 1:l_k
    for i = 1:l_h

        if h(i) > 1 & k(j) > 1
            flD(i,j) = 1;
        elseif h(i) < -1 | k(j) < -1
            flD(i,j) = 0;
        elseif h(i) > 1 & k(j) < 1
            h(i)=1;
            n=(1+h(i))/0.05;
            m=(1+k(j))/0.05;
            n=round(n);
            m=round(m);
            flD(i,j) =fluxXX( n+1,m+1);
        elseif h(i) < 1 & k(j) > 1
            k(j)=1;
            n=(1+h(i))/0.05;

```

```

        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flD(i,j) =fluxXX( n+1,m+1);

    else

        n=(1+h(i))/0.05;
        m=(1+k(j))/0.05;
        n=round(n);
        m=round(m);
        flD(i,j) =fluxXX( n+1,m+1);

    end

end

end

fl=(flA+flC-flB-flD);
A_i = sum(fl);
flu_i = sum(A_i);
energy = flu_i*(0.05*Rs)*(0.05*Rs);
factor = Area_i/energy;
if quandant == 2
    if Xo > -0.15 | Yo < 0.15
        fl = 0*fl;
    else
        fl= fl;
    end
elseif quandant == 4
    if Xo < 0.15 | Yo > -0.15

```

```

    fl = 0*fl;

else

    fl= fl;

end

else

fl= fl;

end

h12=[Ax-6*Rs:0.05:Bx+6*Rs];
k12=[Ay-6*Rs:0.05:Cy+6*Rs];

l_h=length(h12);
l_k=length(k12);

km=[ cos(theta_12) cos(sb);
    sin(theta_11) sin(sa)];

Euler_INV=inv(Euler);

cylindrical=zeros(s_height,s_cet);

% transform image plane to x-y-z
for j=1:l_k
    for i=1:l_h
        xx(i,j)=(h12(i)*km(1,1)+k12(j)*km(1,2))*Rs;
        yy(i,j)=(k12(j)*km(2,2)+h12(i)*km(2,1))*Rs;
        Xx(i,j)=xx(i,j)*Euler_INV(1,1)+yy(i,j)* Euler_INV (1,2)+X0p;
        Yy(i,j)=xx(i,j)* Euler_INV (2,1)+yy(i,j)* Euler_INV (2,2)+Y0p;
        Zz(i,j)=xx(i,j)* Euler_INV (3,1)+yy(i,j)* Euler_INV (3,2)+Z0p;
        Xx(i,j)=Xx(i,j)*100;
        Yy(i,j)=Yy(i,j)*100;
        Zz(i,j)=round(Zz(i,j)*100);
    end
end

```

```

rho1(i,j)=atan2(Yy(i,j),Xx(i,j));
rho1(i,j)=rho1(i,j)*180/pi;
if rho1(i,j) >= 0
    rho1(i,j)= rho1(i,j);
else
    rho1(i,j)= 360+rho1(i,j);
end
rho1(i,j)=round(rho1(i,j));
tt(i,j)=(sqrt(Xx(i,j)^2+Yy(i,j)^2)-200)/sin(tower_al);
X(i,j)=Xx(i,j)+tt(i,j)*Rop(1,1);
Y(i,j)=Yy(i,j)+tt(i,j)*Rop(1,2);
Z(i,j)=Zz(i,j)+tt(i,j)*Rop(1,3);
Z(i,j)=round(Z(i,j));
rho2(i,j)=atan2( Y(i,j),X(i,j));
rho2(i,j)=rho2(i,j)* 180/pi;
if rho2(i,j) >= 0
    rho2(i,j)= rho2(i,j);
else
    rho2(i,j)= 360+rho2(i,j);
end
rho2(i,j)=round(rho2(i,j));
if fl(i,j) > 1
    fl(i,j) == 1;
else
    fl(i,j) = fl(i,j);
end
cylindrical(Z(i,j),rho2(i,j)+1)=fl(i,j)*sin(tower_al)* factor;

```

```
end
end
Flux = cylindrical + Flux;
end
end
xlabel('theta(degree)')
ylabel('tower height (cm.)')
zlabel('flux dimensionless / Fo')
title('solar flux density distribution of 12/1/1')
surf(Flux(900:1100,1:360))
view([-27,66])
```

APPENDIX B

RELATION OF SOLAR INTENSITY AND SOLAR BEAM IRRADIANCE

The solar intensity is related to the solar beam irradiance at normal incidence G_{bn} . From equation (3.12) we can draw the relation between $S_o (W.m^{-2}.sr^{-1})$ and the normal flux density $G_{bn} (W.m^{-2})$ at the consider time.

$$G_{bn} = \int_0^{4\pi} S(\alpha) d\Omega$$

where Ω is solid angle

Solid angle is defined as the ration of the area ds of a spherical surface to the square of its radius d , it can be written as follows

$$d\Omega = \frac{ds}{d^2}$$

$$\int_0^{R_s} d\Omega = \int_0^{R_s} \frac{d(\pi R^2)}{d^2}$$

$$\Omega = \int_0^{R_s} \frac{2\pi R d(R)}{d^2}$$

$$\Omega = \int_0^{R_s} \frac{2\pi (d \tan \alpha) d(d \tan \alpha)}{d^2}$$

$$\Omega = \int_0^{\alpha_s} \frac{2\pi d^2 \tan \alpha \sec^2 \alpha d\alpha}{d^2}$$

$$\Omega = \int_0^{\alpha_s} 2\pi \tan \alpha \sec^2 \alpha d\alpha$$

then

$$G_{hm} = \int_0^{\alpha_1} S(\alpha) \cdot 2\pi \tan(\alpha)(1 + \tan^2 \alpha) d\alpha$$

REFERENCES

- [1] World Energy Council. Measuring Solar Insolation [Online]. Available from [http://www. Worldenergy.org/wecgeis/publications/reports/ser](http://www.Worldenergy.org/wecgeis/publications/reports/ser) [2005, Mar 2].
- [2] DNT, NESDB, and oil traders in pursuance of section 7. Thailand Energy Situation 2003 [Online]. Available from <http://www.dedg.go.th> [2005, Mar 2].
- [3] Solar PACES. Solar Thermal Power Plants [Online]. Available - <http://www.solarpaces.org/index.html> [2005, Mar 2].
- [4] เสริม จันทรฉาย และคณะ. แผนที่ศักยภาพพลังงานแสงอาทิตย์จากข้อมูลดาวเทียมสำหรับประเทศไทย. ใน แผนที่ศักยภาพพลังงานแสงอาทิตย์เฉลี่ยต่อปี , หน้า 82 กรุงเทพมหานคร: จีรังริชต์, 2542.
- [5] Moustafa M. Elsayed and Kadry A. Fathalah. Estimation of percentage useful area of a heliostat when considering shadowing and blocking. Solar & Wind Technology **3** (1986): 199-205.
- [6] Lipps, F.W., Vant-Hull, L.L. A cell wise method for the optimization of large central receiver systems. Solar Energy **20** (1978): 505-516.
- [7] F.M.F. Siala, M.E. Elayeb. Mathematical formulation of a graphical method for a no-blocking heliostat field layout. Renewable Energy **23** (2001): 77-92.
- [8] S. Puliaev, J.L. Penna, E.G. Jilinski, and A.H. Andrei. Solar diameter observation at Observatorio Nacional in 1998-1999. Astron. Astrophys **143** (2006): 265-267.

- [9] M.D. Walzel, F.W. Lipps and L.L. Vant-Hull. A solar flux density calculation for a solar tower concentrator using a two-dimensional Hermite function expansion. Solar Energy **19** (1977): 239-253.
- [10] G.A. Smith. "Monte-Carlo ray trace simulation for solar central receiver system" UC-62, under ERDA Contract AT (29-1)-789, University of Houston. 10-11 August 1977.
- [11] F.W. Lipps and M.D. Walzel. An analytic evaluation of the flux density due to sunlight reflected from a flat mirror having a polygonal boundary. Solar Energy **21** (1978): 113.
- [12] G.L. Schrenk. "The role of simulation in the development of solar-thermal energy conversion systems" Proc. 11th Intersociety Energy Conversion Engineering Conf. 12-17 September 1976.
- [13] M.M.Elsayed and K.A. Fathalah. Solar flux density distribution using a separation of variable/superposition technique. Renewable energy **4** (1994): 77-87.
- [14] J. C. Hennem and J.L. Abatut. An analytical method for reflected flux density calculation. Solar Energy **32**(1984): 357-363.
- [15] สมชาย เกียรติกมลชัย และคณะ. รายงานฉบับสมบูรณ์ โครงการศึกษา ออกแบบและสร้างต้นแบบเตาเผาสุริยะอุณหภูมิสูง. ใน การวางตำแหน่งฮีลิโอสแตต. หน้า 100-106 กรุงเทพมหานคร: ฝ่ายวิชาการ จุฬาลงกรณ์มหาวิทยาลัย, 2547.
- [16] M.R.Riaz. A Theory of Concentrators of Solar Energy on a Central Receiver for Electric Power Generation. Transactions of ASME **32**(1976): 375-386.

VITAGE

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Conference Presentations:

1. Aparporn Sakulkalavek and Somchai Kiatgamolchai, "Design of Heliostat field for a Central Receiver System and Solar Flux Energy Calculation", *The 4th Conference on Science and Technology*, Thammasat University, Pathumthani, Thailand, March 16 (2005).
2. Aparporn Sakulkalavek and Somchai Kiatgamolchai, "A Calculation of Solar Flux Distribution on a Cylindrical Receiver Surface of a Central Receiver System", *31st Congress on Science and Technology of Thailand (STT 2005)*, Suranaree University of Technology, Nakhon Ratchasima, Thailand, October 18-20, (2005).

